2DBAT to the separated circular motions. The rotation curve reflecting the total kinematics of the LMC, dark and baryonic matters is then be combined with the mass models of baryons, mainly stellar and gaseous components in order to examine the dark matter distribution. Here, we present the analysis of the extracted HI gas maps, rotation curve, and J, H and K-band surface photometry of the LMC.

[포 GC-04] HI gas kinematics of galaxy pairs in the Hydra cluster from ASKAP pilot observations

Shin-Jeong Kim¹, Se-Heon Oh², and ASKAP WALLABY Science Working Group2 (SWG2)

¹Department of Astronomy and Space Science, Sejong University, Seoul, Korea

²Department of Physics and Astronomy, Sejong University, Seoul, Korea

examine the HI gas kinematics and distribution of galaxy pairs in group or cluster environment from high-resolution Australian Kilometre Array Pathfinder (ASKAP) observations. pilot We well-resolved galaxies in the Hydra cluster of which 4 galaxies are visually identified as pairs and others are isolated ones. We perform profile decomposition of HI velocity profiles of the galaxies using a new tool, BAYGAUD which enables us to separate a line-of-sight velocity profile into an optimal number of Gaussian components based on Bayesian MCMC techniques. All the HI velocity profiles of the galaxies are decomposed into kinematically cold or warm gas components with their velocity dispersion, 4~8 km/s or > 8 km/s. respectively. We derive the mass fraction of the kinematically cold gas with respect to the total HI gas mass, $f = log10(M_cold / M_HI)$, of the galaxies and correlate them with their dynamical mass. The cold gas reservoir of the paired galaxies in the Hydra cluster is found to be relatively higher than that of the isolated ones which show a negative correlation with the dynamical mass in general.

[포 GC-05] Galaxy overdensity around sub-mm sources from SPT-SZ survey

Yeonsik KIM¹, Hyujin Shim²
¹Department of Astronomy and Atmospheric Sciences, Kyungpook National University,
²Department of Earth Science Education, Kyungpook National University

We study the overdensity of near-infrared sources around 508 sub-mm sources classified as

dusty galaxies in the SPT-SZ survey catalog observed in 95 GHz (3.15 mm), 150 GHz (2 mm) and 250 GHz (1.2 mm) bands. We used the VISTA hemisphere survey data release 6 (VHS DR6) catalog covering the J, H, Ks bands. The mean number of galaxies within a radius of 60 arcsec (corresponding to about 500 kpc at z=2) from 500 randomly selected positions is 14.4, while the galaxy number distribution is approximated as a Gaussian with a standard deviation of 7.9. From the 2500 deg2 of SPT-SZ survey + VHS DR6 data, there were 27 sub-mm sources that have galaxy overdensity higher than 4σ . We color-magnitude diagram around 27 sub-mm sources with enhanced galaxy surface densities, in order to investigate the presence of structure around sub-mm sources.

[王 GC-06] GAS KINEMATICS AND PHOTOIONIZATION IN TYPE 1 AGNs WITH STRONG OUTFLOWS

CHANGSEOK KIM^1 , JONG-HAK $\mathrm{WOO}^1,\ \mathrm{RONGXIN}$ LUO^2

¹Astronomy Program, Department of physics and Astronomy, Seoul National University 151-742, Korea

²Shanghai Astronomical Observatory, 80 Nandan Road, Shanghai 200030, China

We present spatially resolved outflows and photoionization for a pilot sample of 11 type 1 AGNs (z<0.3) based on the Gemini Multi-Object Spectrograph Integral Field Unit data. These AGNs were selected since we found strong outflow signatures in SDSS spectra. We focus on [OIII] and Hα emission lines to probe outflow kinematics by measuring line flux, velocity, and velocity We dispersion at each pixel. investigate characteristics of gas kinematics of type 1 AGNs and compare them with those of type 2 AGNs in our previous studies. Furthermore, by drawing BPT map, photoionization states will be also discussed. Based on the results, we discuss various implications on the impacts of outflows on star formation in host galaxies.

[포 GC-07] Gas kinematics and star formation in NGC 6822

Hye-Jin Park¹, Se-Heon Oh², Jing Wang^{3,4}, Yun Zheng^{3,4}, Hong-Xin Zhang^{5,6}, and W.J.G. de Blok^{7,8,9}
¹Department of Astronomy and Space Science, Sejong University, Seoul, Korea
²Department of Physics and Astronomy, Sejong University, Seoul, Korea
³Kavli Institute for Astronomy and Astrophysics