측광 자료에 대하여 Period04를 활용한 power spectrum 분석을 통해 변광 주기를 찾았다. 이 분석을 통해 2개의 변광 주기 f_1 =7.127c/d, f_2 =1.995c/d 를 찾았다. 2개의 주기 가운데 f_1 은 fundamental mode로서 Gregor(2009)가 처음으로 제시한 주기 0.1389 day와 일치하는 것이며, f_2 는 이 연구에서 새롭게 찾은 주기로서 non-radial mode에 해당되는 것으로 해석하였다. 분석 결과를 바탕으로 HD5371의 변광특성과 물리적 성질을 논하였다. HD5371의 완전한 주기 해를 얻기 위한 장기 관측을 수행 중에 있다.

[\(\mathbb{X}\) ST-06] Comparison of multi-planetary systems including hot-Super Earth with and without exo-Jupiter

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Almost hot-Super Earths (Rp~1 to 4Rearth orbital period < 100 days) are around Sun-like stars. But our solar system does not have hot-Super Earth. Andre et al. 2015 has explained this phenomenon by that Jupiter blocks migration of super earth. We have found a multi-planetary system KOI-94 with exo-Jupiter and hot-Super Earth from NASA exoplanet archive data (http://exoplanetarchive .ipac.caltech.edu). In this study, within multi-planetary systems including hot-Super Earth, we compared those with and without exo-Jupiter using their host star and exoplanet parameters, such as metallicity [Fe/H], T_{eff} and R_{*}/R_p.

[至 ST-07] O VI Raman spectroscopy of the S-type symbiotic star V455 Sco

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We present the high-resolution spectrum of the S-type symbiotic star V455 Sco obtained with the Dupont telescope in 2014 June. We note that the Raman-scattered O VI λ 1032 at 6825 Å exhibits a triple-peak profile. Adopting an accretion disk model with an additional contribution from a collimated bipolar outflow, we attempt to fit the profile. We propose that the blue and central peaks are formed via Raman-scattering of O VI line photons from the accretion flow and that the bipolar flow is responsible for the remaining red peak. It is also noted that V455 Sco exhibits the Raman-scattered He II features blueward of H α

and H B.